


ADVANCING THE TECHNICAL READINESS OF THE MOMA MINIATURE LINEAR ION TRAP MASS SPECTROMETER.

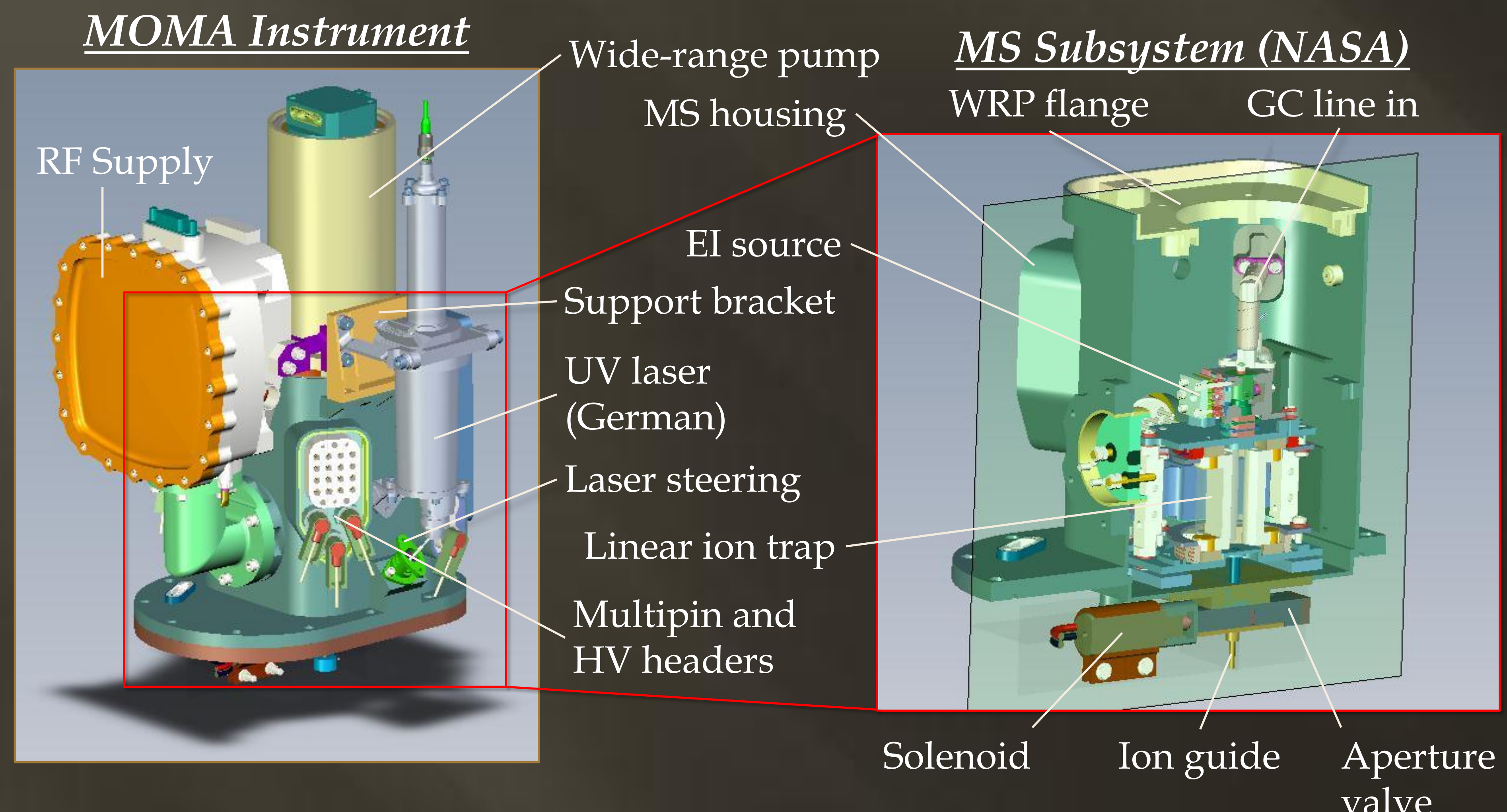
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INTRODUCTION

The primary science goal of the joint ESA-Roscosmos-NASA ExoMars Program is to search for and identify signs of past or present life on Mars through a chemical investigation of the martian atmosphere (via an orbiter to be launched in 2016) and surface (via a rover to be launched in 2018). The Mars Organic Molecule Analyzer (MOMA), a dual-source, mass spectrometer-based instrument capable of both pyrolysis-gas chromatography (pyr-GC) and laser desorption/ionization (LDI), is a key instrument on the Pasteur Payload of the ExoMars 2018 rover. When combined with the unprecedented two-meter depth sampling capability of the current rover, MOMA affords a uniquely broad and powerful search for organics over a range of preservational environments, volatility, and molecular weight. In addition to enabling mission science, MOMA critically informs strategies for both sampling and *in situ* analysis for Mars Sample Return (MSR).

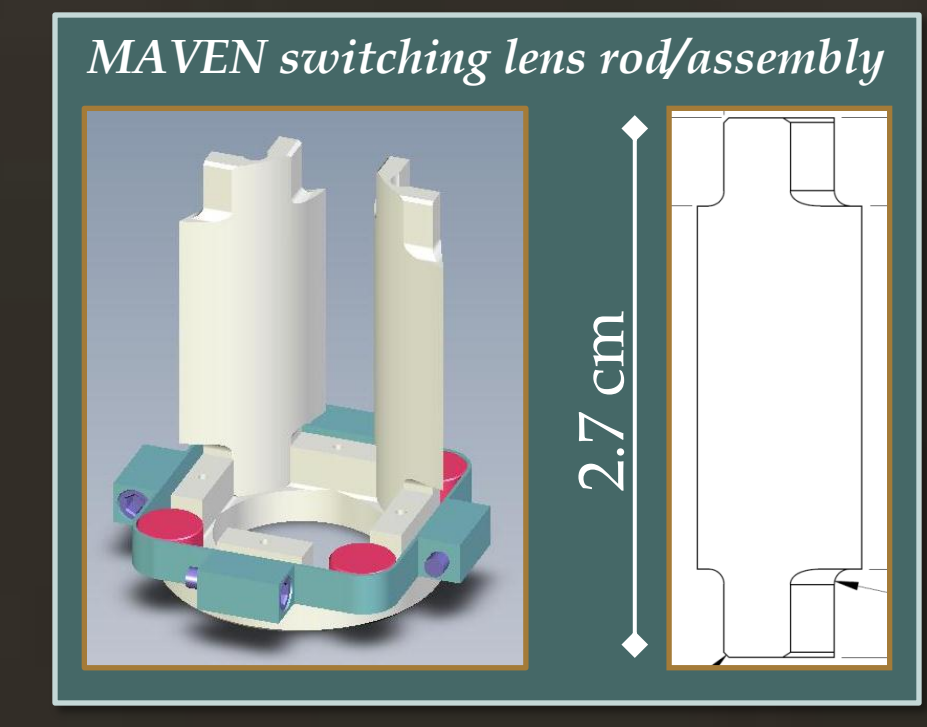
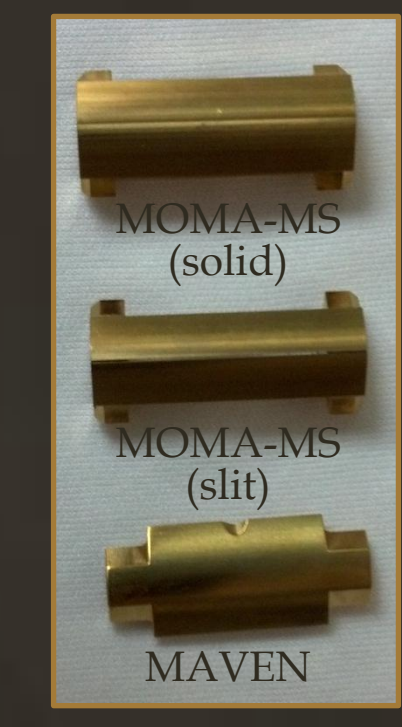
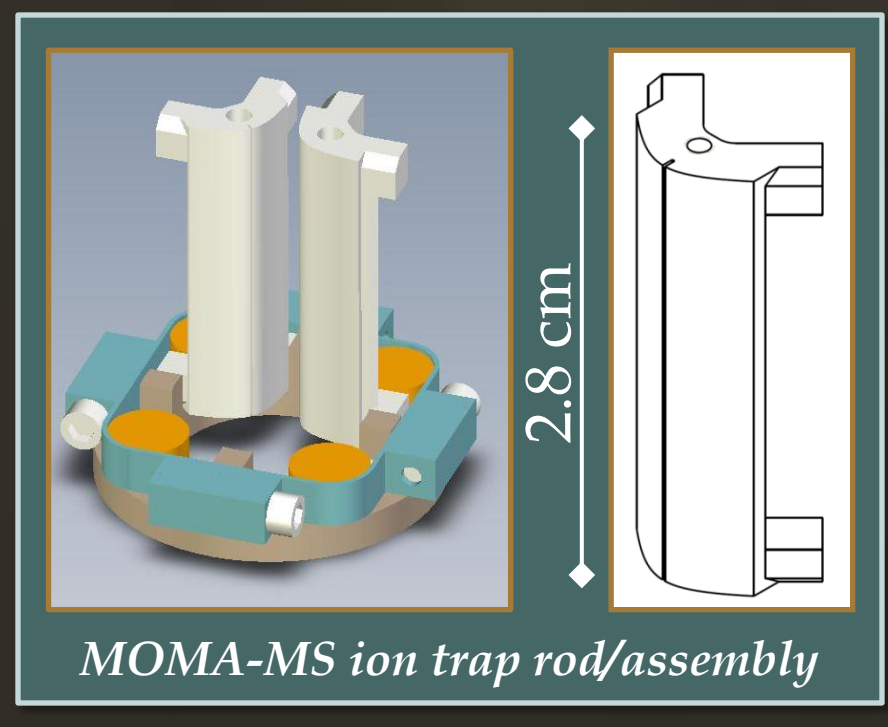
The NASA contribution to the MOMA instrument comprises the mass spectrometer (MS) subsystem, including the following components shown below.



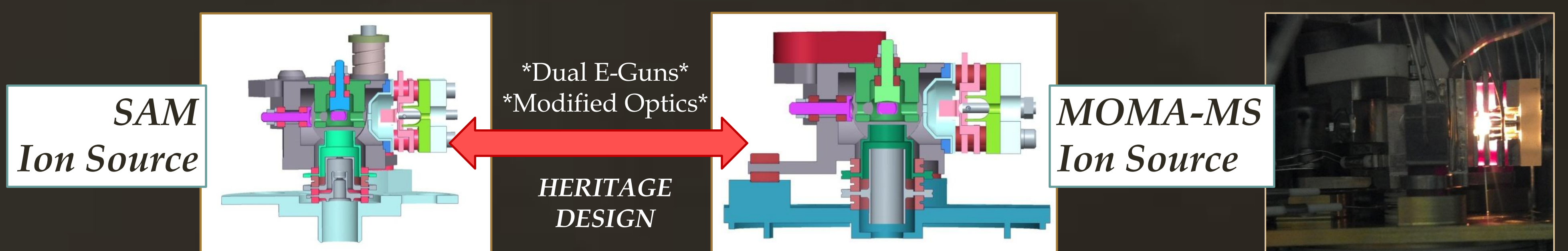
HERITAGE (HIGH-MATURITY) HARDWARE COMPONENTS

(Based on SAM/LADEE/MAVEN designs)

1. RF Supply, Housing, and Electrical Headers/Feedthroughs
2. Create 100k RPM Wide-Range Pump
3. Electron Ionization Source
4. Hyperbolic Rods

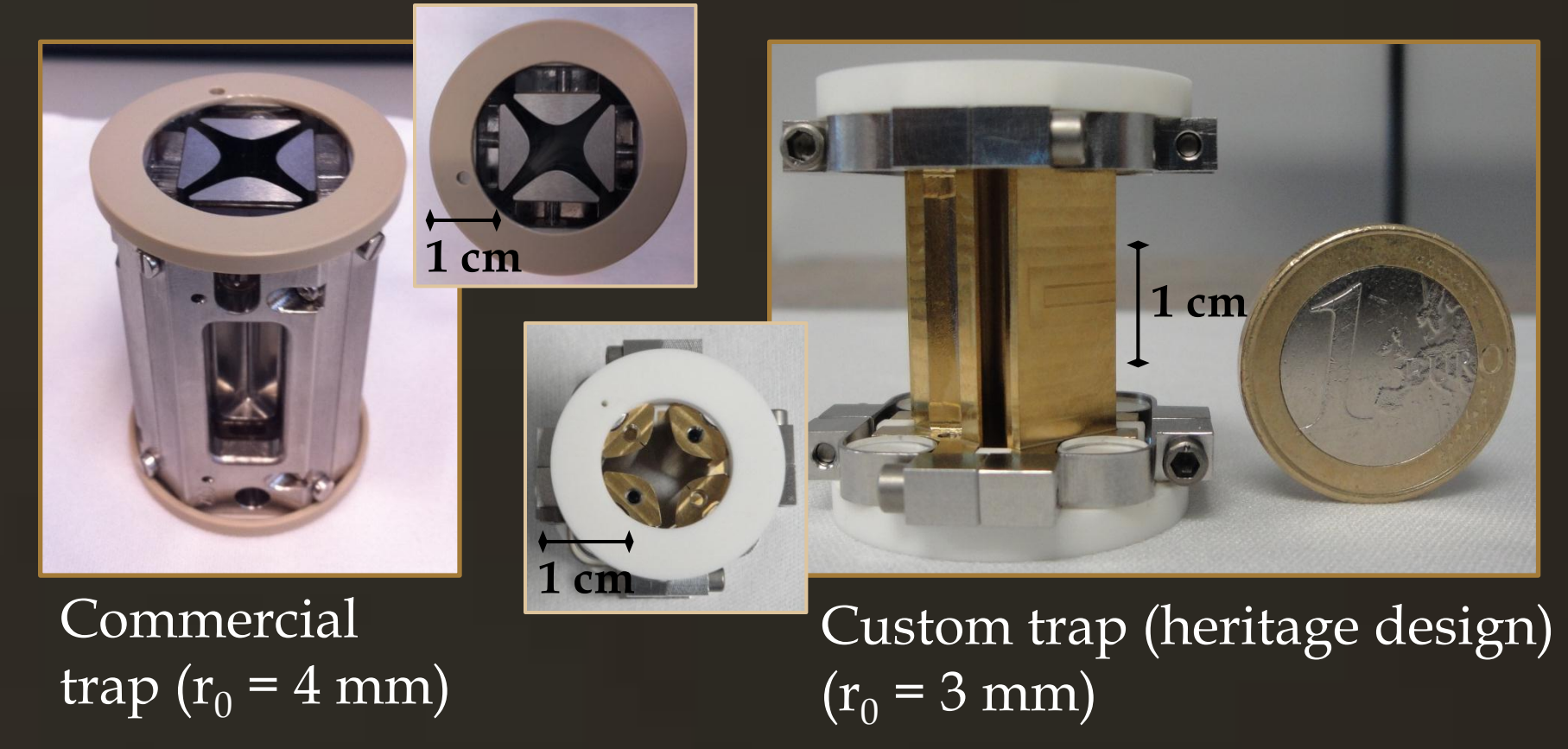


SAM vs. MOMA-MS 100k rpm WRPs
 (NOTE: this is a backup option for MOMA-MS, as the baseline is currently a 200k rpm pump that requires qualification)



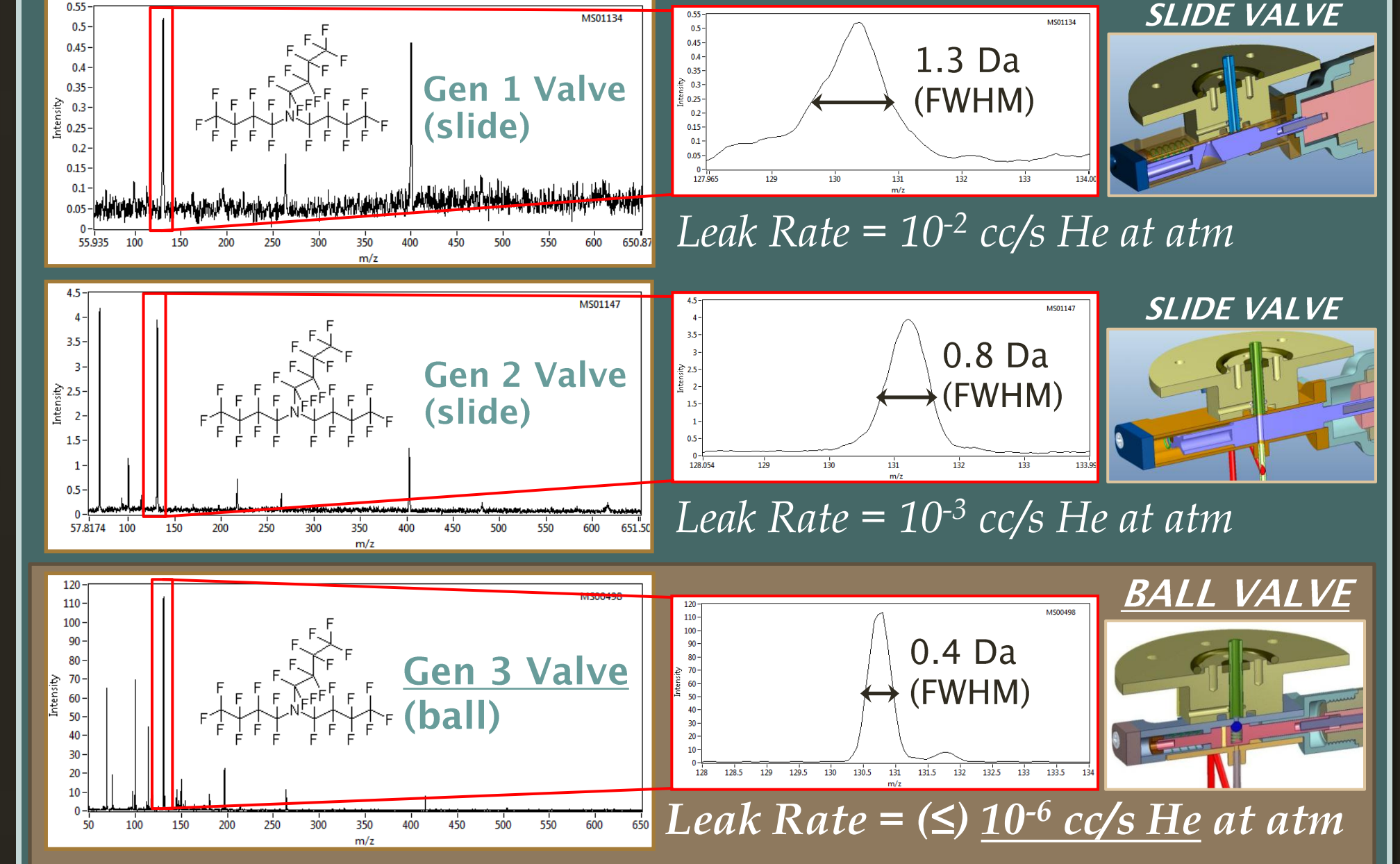
NEW TECHNOLOGY: LINEAR ION TRAP MASS SPECTROMETER

The heart of MOMA-MS is a linear (or 2-D) ion trap mass analyzer. Compared to 3-D varieties, linear ion traps are particularly suited for FLT as they provide: a symmetrical interface to two distinct ion sources (e.g., pyr-GC and LDI modes); increased dynamic range; improved ion trapping capacity and efficiency; redundant detector/dynode assemblies; and, an analyzer design based on a set of four hyperbolic rod electrodes, similar to GSFC heritage quadrupole mass filters/lenses (SAM/LADEE/MAVEN).



Laser desorption/ionization (LDI) mass spectrometry with the MOMA-MS linear ion trap is supported at Mars ambient pressures by the implementation of an LDI inlet consisting of an aperture valve and ion guide.

Low Maturity Component: Aperture Valve



Evolution of MS "Guts"
GB1: Breadboard demonstrated proof-of-concept and allowed evaluation of trade space for system performance

- Commercial $r_0 = 4$ mm ion trap
- Shielded single-gun EI source
- Competing CEM detectors
- Shielded rod assembly

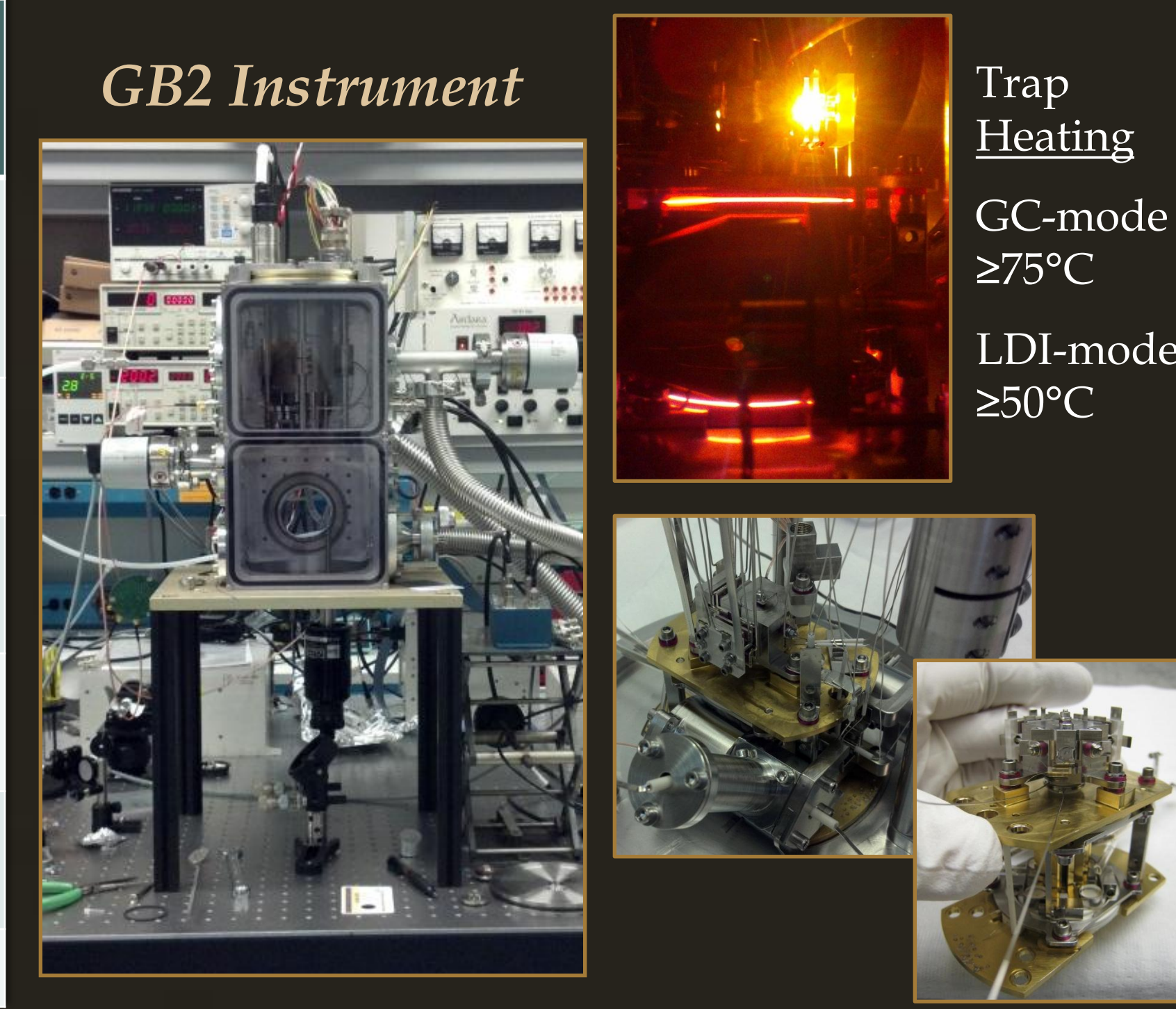
GB2: FLT-like instrument dedicated to demonstrating TRL-6 performance requirements under environmental conditions

- Custom miniaturized $r_0 = 3$ mm ion trap
- Shielded detector/dynode assemblies
- Brassboard RF and DET electronics
- Heated ion trap and EI lenses
- Dual gun EI source

MOMA-MS TRL-6 PERFORMANCE REQUIREMENTS

Across a defined range of martian surface conditions (i.e., P = 5 to 7 torr, T = -40°C to +20°C), the integrated MOMA-MS subsystem (including constituent high- and low-maturity components) must meet the following performance specifications to ensure the full science return over the 180 sol operational lifetime of the mission:

Performance Specification	MOMA-MS pyr-GC mode	MOMA-MS LDI mode
Sensitivity	≤ 1 nmol (SNR ≥ 10) of PFTBA	≤ 1 pmol (SNR ≥ 3) of Rhodamine 6G and Angiotensin II
Mass Range	50 - 500 amu	50 - 1000 amu (Goal: up to 2000 amu)
Resolution (FWHM)	≤ 1 amu over mass range	≤ 1 amu (50 - 500 amu) ≤ 2 amu (501 - 1000 amu)
Analytical Drift	< 0.4 amu per experiment	< 0.4 amu per experiment
Mass Accuracy	< 0.4 amu over mass range	< 0.4 amu over mass range
Dynam Range	≥ 100	≥ 100



These performance metrics have been demonstrated on the GB2 prototype instrument, under FLT-like conditions and with active heating of the ion trap mass analyzer (see Pinnick et al. poster).

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