

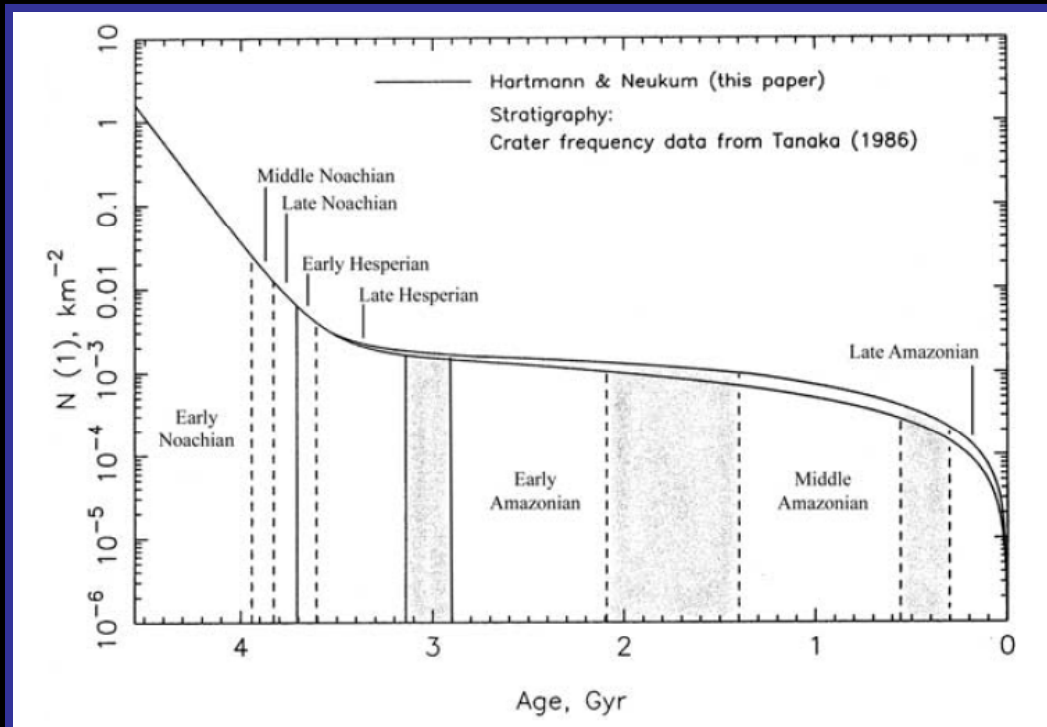
A New Approach to In-Situ K-Ar Geochronology

Better than 10% precision on K/Ar ages with a single analytical instrument, easily accessible Ar extraction temperatures, and no mass determination required

J.A. Hurowitz, K.A. Farley, N.S. Jacobson,
P.D. Asimow, J.A. Cartwright*

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Crater Counting Chronology

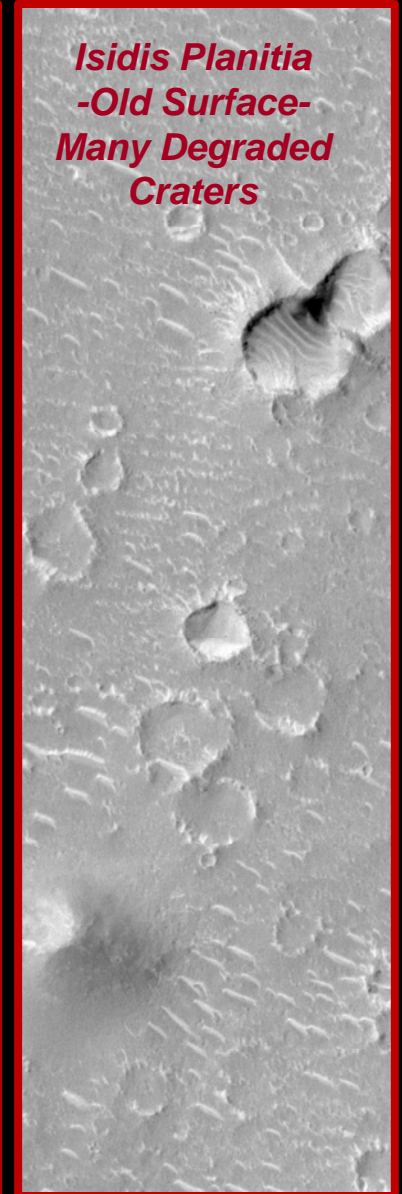
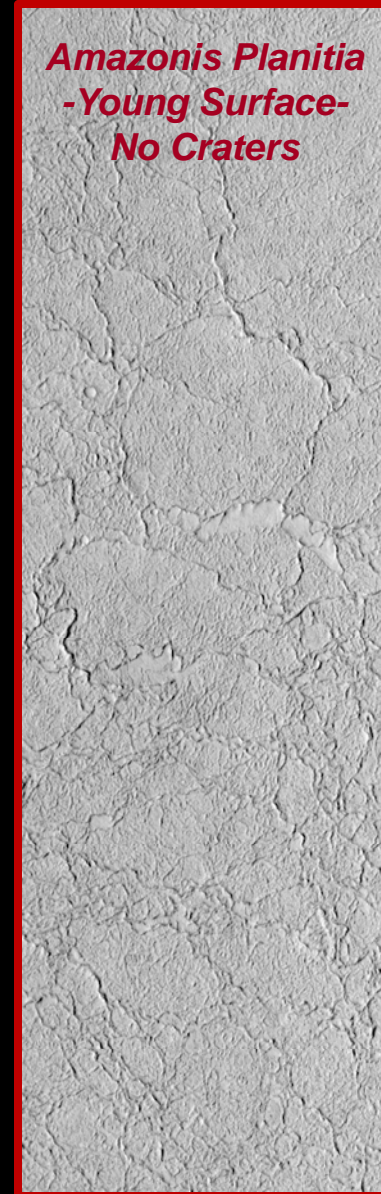


The Two Major Assumptions (weaknesses):

1. The lunar cratering rate can be scaled to an appropriate value for Mars
2. The geologically active surface of Mars (which erases and modifies craters in ways not experienced on the Moon) can be properly accounted for

Amazonis Planitia
-Young Surface-
No Craters

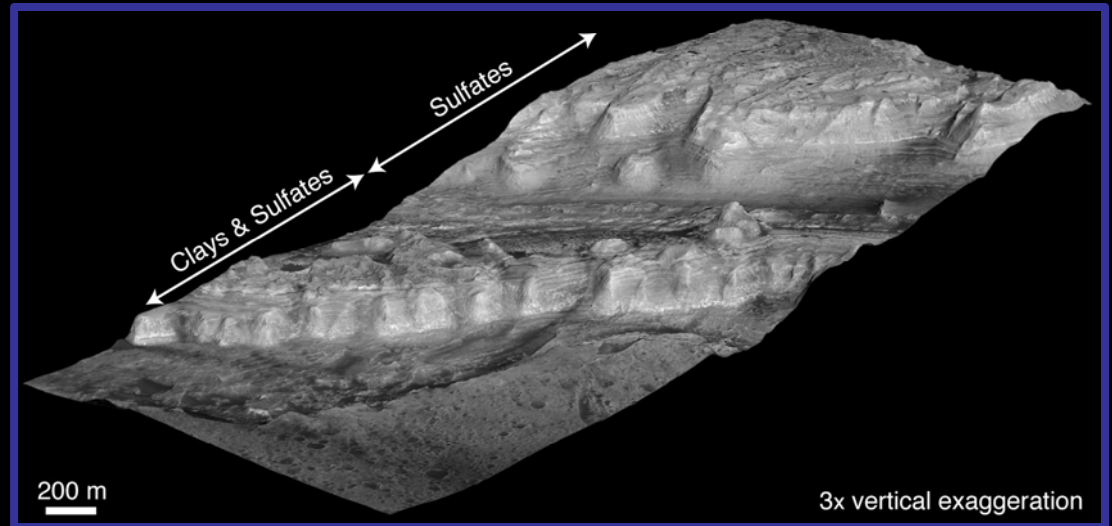
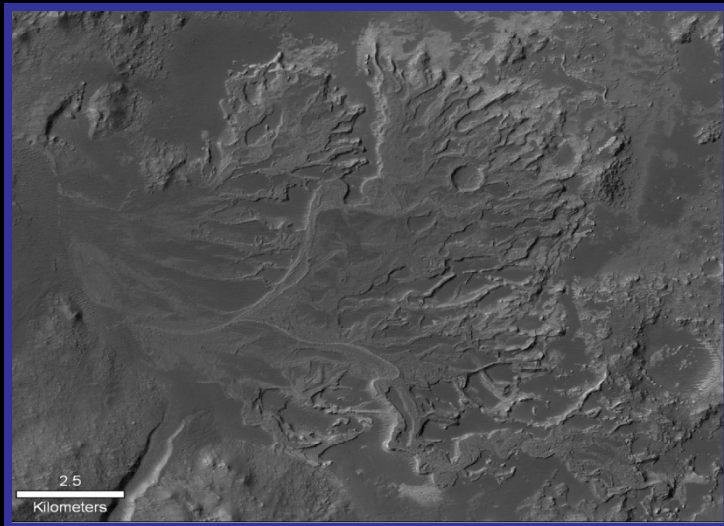
Isidis Planitia
-Old Surface-
Many Degraded
Craters



Why in-situ Geochronology?

Important Questions:

- 1) *When did aqueous activity occur on the Martian surface?*
- 2) *When did the Martian surface environment undergo major changes?*
- 3) *Does Mars record evidence of a late heavy bombardment?*
- 4) *Did Mars harbor a biosphere, and if so, exactly when?*



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These are fundamentally questions of timing. Our understanding of absolute time on Mars is currently subject to the limitations of crater counting.

K-Ar Geochronology

What?

- $^{40}\text{K} \rightarrow ^{40}\text{Ar}$, $t_{1/2} = 1.3 \times 10^9$ years

Why?

- Applicable to basalt and K-bearing alteration phases including jarosite and alunite
- Useful over a wide range of age (10's of Ma to many Ga)
- Involves K, a major element (rather than trace elements, e.g., Rb-Sr, Sm-Nd, U-Pb)
- Martian targets of interest are likely to be old, and accordingly, have high Ar-concentration

How are K-Ar Ages Measured on Earth?

1. Measure K-concentration by conventional techniques
2. Place a second **weighed** aliquot of sample into an MS
3. Heat sample to **melting point** ($\sim 1200^{\circ}\text{C}^+$ for basalts)
4. Measure isotopic composition and amount of Ar
5. Calculate K-Ar age using K/Ar concentration ratio

Weighing small aliquots of sample
+
Heating them above 1000°C
=
Some of the hardest things to do with
spacecraft instrumentation.

How do we surmount these problems?

*A work around to designing a space-qualified high temperature oven
= FLUX ASSISTED MELTING*

TABLE 3

Common fluxes used in sample decomposition by fusion ¹

| Flux | Melting point °C |
|---|---------------------------|
| Lithium metaborate, LiBO_2 | 845 |
| Lithium tetraborate, $\text{Li}_2\text{B}_4\text{O}_7$ | 930 |
| Sodium peroxide, Na_2O_2 | 480 (decomposed) |
| Sodium carbonate, Na_2CO_3 | 851 |
| Sodium hydroxide, NaOH | 318 |
| Potassium carbonate, K_2CO_3 | 891 |
| Potassium hydroxide, KOH | 360 |
| Potassium pyrosulfate, $\text{K}_2\text{S}_2\text{O}_7$ | 419 |
| Sodium pyrosulfate, $\text{Na}_2\text{S}_2\text{O}_7$ | 403 |
| Ammonium iodide, NH_4I | > 300 (starts to sublime) |

¹ Data from Potts (1987), Bock (1979), and Erdey et al. (1964).



SAM Pyrolysis oven can achieve 1000-1100°C

How do we surmount these problems?

*A work around to designing a space-qualified
high precision analytical balance
= DOUBLE ISOTOPE DILUTION*



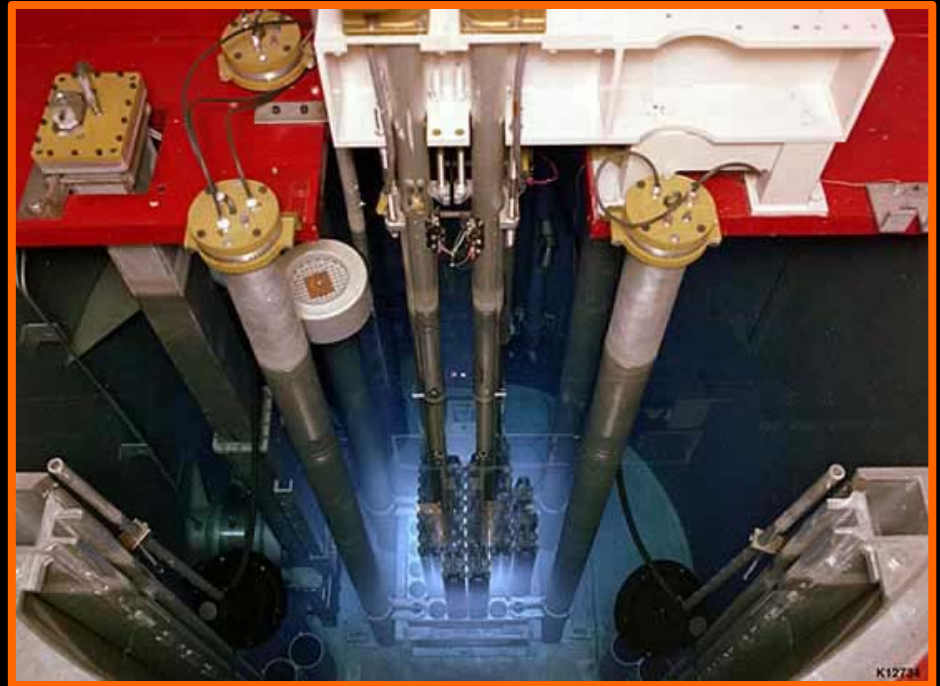
- Isotope dilution is the “gold standard” for laboratory measurements, yielding typical precisions of ~1%
- We employ a double-isotope spike containing ^{41}K and ^{39}Ar
- Here’s how we cancel out the mass measurement:

$$\frac{{}^{40}\text{Ar}^*}{{}^{40}\text{K}} = C * \frac{\left(\frac{{}^{40}\text{Ar}^*}{{}^{39}\text{Ar}}\right)_{\text{measured}}}{\left(\frac{{}^{39}\text{K}}{{}^{41}\text{K}}\right)_{\text{measured}}} * \left(\frac{{}^{41}\text{K}}{{}^{39}\text{Ar}}\right)_{\text{splke}}$$

Spike Synthesis



Caltech Petrology Lab

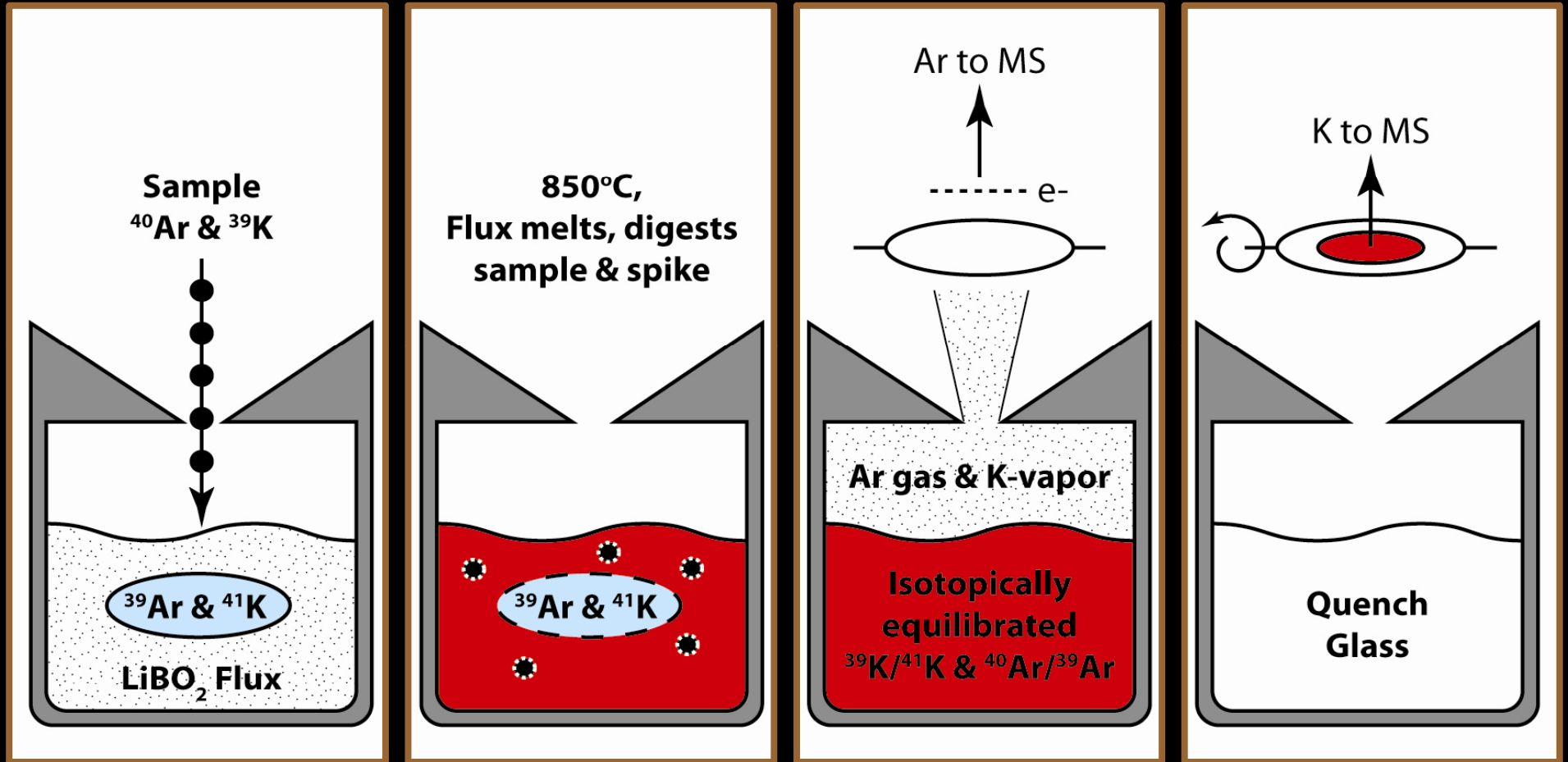


Oregon State University's TRIGA Reactor

Synthesized a glass containing known amounts of ^{41}K & ^{39}Ar :

- We did this by combining synthetic albite glass ($\text{NaAlSi}_3\text{O}_8$) with isotopically-enriched ^{41}KCl and melting them in a controlled atmosphere.
- This glass was irradiated, producing ^{39}Ar from the trace ($\sim 1\%$) ^{39}K present in the isotopically enriched ^{41}KCl .

How might we employ our technique on another planet?



$$\frac{^{40}\text{Ar}^*}{^{40}\text{K}} = C * \frac{\left(\frac{^{40}\text{Ar}^*}{^{39}\text{Ar}}\right)_{\text{measured}}}{\left(\frac{^{39}\text{K}}{^{41}\text{K}}\right)_{\text{measured}}} * \left(\frac{^{41}\text{K}}{^{39}\text{Ar}}\right)_{\text{splike}}$$

Our First Age Measurement

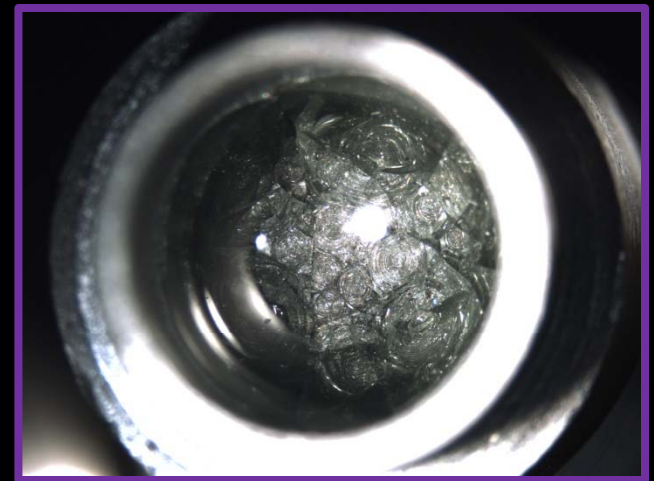
At present = no single instrument for both Ar & K measurement...

→ We are using 2 steps to perform this analysis using a basalt sample from the Viluy traps, Siberia (0.8 wt% K_2O).

AIM: Measure an age of 354 ± 2 Ma

Step 1:

Combine flux (150mg) + basalt (15mg) + spike glass (1.5mg), in a crucible. Melt at 950°C , measure Ar-isotopic composition on Ken Farley's noble gas mass spectrometer at Caltech.

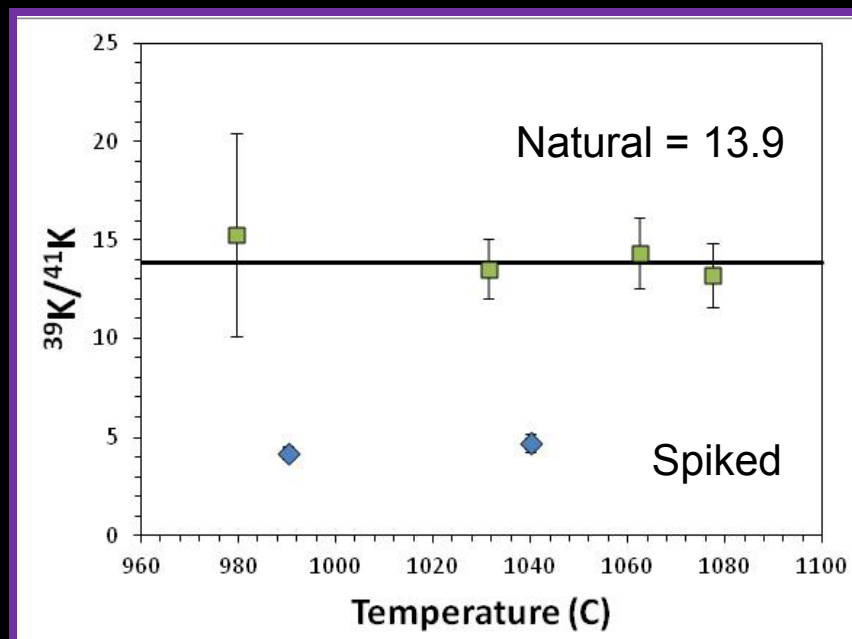
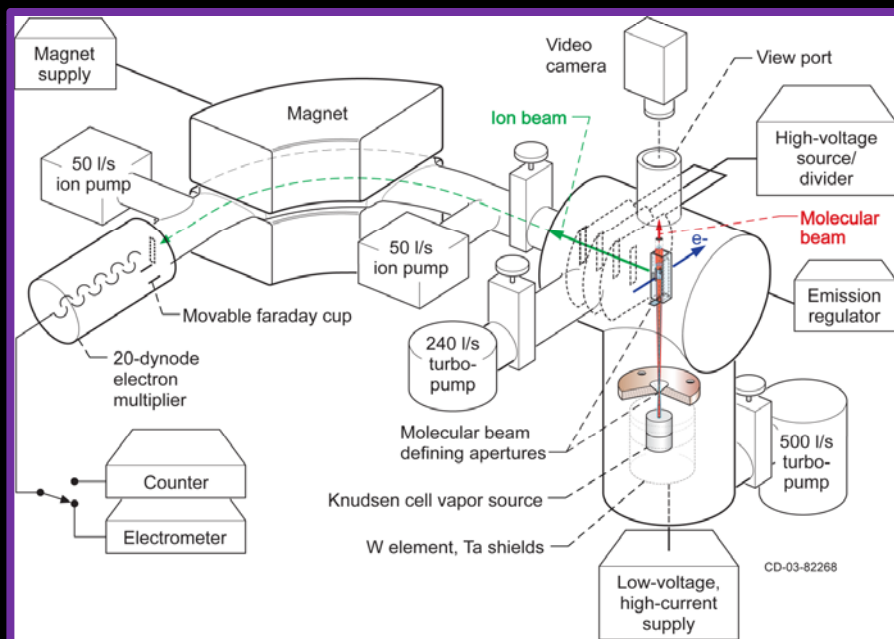


Our First Age Measurement

Step 2:

FedEx the cell to NASA Glenn Research Center, measure K isotopic composition on Nate Jacobson's Knudsen Effusion Mass Spectrometer.

AIM: Measure an age of 354 ± 2 Ma



RESULT: We calculate a preliminary age of 337 ± 30 Ma

Conclusions

1. Developed and verified a technique that can be implemented for a flight instrument system. Basically consists of an oven and a mass spectrometer
→ Uses a mixture of basic geochemical practices developed in the mid-20th century, including K-Ar, flux digestion & isotope dilution
2. No mass measurement is required, high temperatures are not required, MS only has to measure isotope ratios, and no new technology development is required.
3. This double isotope dilution technique should yield in-situ whole rock ages with precision better than 10%.